

1 **MINERAL PHYSICS OF EARTH CORE: IRON ALLOYS**
2 **AT EXTREME CONDITION**

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12 **Abstract** Mineral physics and high-pressure crystallography research
13 provide crucial information on the enigmatic properties of the Earth's interior
14 and is motivating multidisciplinary efforts to re-evaluate our understanding
15 of the past and current states of the planet. We discuss the structures,
16 melting temperatures, sound velocities of the dominant core-forming iron-
17 nickel-light element alloys, as well as seismic anisotropy, thermal and
18 chemical states of the Earth core in view of structural studies of iron-based
19 alloys at multimegabar pressure range.

20 **Keywords:** Earth core, iron alloys, high pressure

21 **1. Introduction**

22 The Earth's core is the most remote region on our planet (Figure 1). The
23 boundary of the core is at about 2,900 km in depth. Not only do we not have
24 samples from the core, we do not even expect to get any. To date, the most
25 direct observations of the core have come from seismological studies using
26 remote-sensing techniques. Due to the complex internal structure of the
27 Earth, seismic investigations require extensive data coverage and appropriate

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28 models. Decoding geochemical signature of the core carried by mantle
 29 plumes faces similar challenges. Experimental and computational simulations
 30 have been hindered by the necessity to approach pressures over 140 GPa
 31 and temperatures above 3,000 K, prevalent in the core. For these reasons,
 32 many fundamental issues concerning the Earth's core remain controversial
 33 and poorly understood.

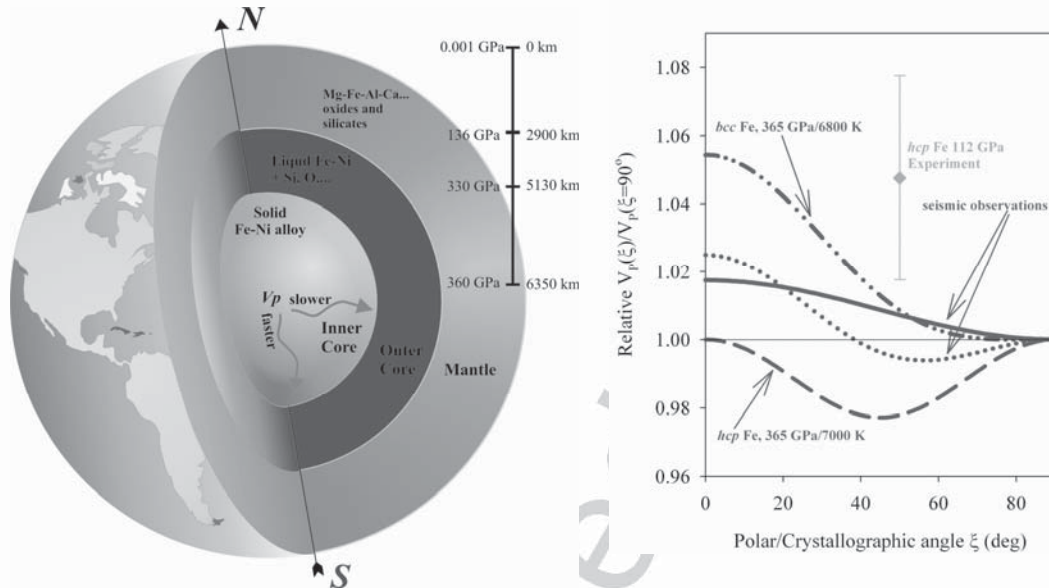


Figure 1. A cutaway of the Earth's interior (left) reveals seismic anisotropy of the inner core (right). Representative elastic anisotropies of single-crystal hcp (dashed line) and bcc (solid line) iron from theory (Belonoshko *et al.*, 2008), prefer-oriented hcp iron from static experiments at 112 GPa (diamond) (Antonangeli *et al.*, 2004), and seismic observations in the inner core are shown for comparison.

34 Despite all these difficulties, the Earth's core is in focus of the highest
 35 research interest due to development of experimental and computational
 36 methods, as well as steady improvement in observational techniques. Recent
 37 studies reveal a number of unusual and enigmatic phenomena related to the
 38 core properties and dynamics. Among them is the discovery of the inner
 39 core anisotropy, *i.e.*, seismic waves travel faster along the Earth's polar axis
 40 than in the equatorial direction. There is also evidence for differential
 41 rotation between the inner core and the rest of the Earth and fine-scale
 42 heterogeneity. Earth's magnetic field strongly links to the core dynamics.
 43 Acceleration of the movement of magnetic poles over the last 150 years is a
 44 clear manifestation of the importance of understanding of the Earth core
 45 properties for human society.

46 2. Composition of the Earth's Core

47 Composition of the Earth's core is a geochemical parameter crucial for
48 understanding the evolution and current dynamics of our planet. Since the
49 discovery of the Earth's core about a century ago, the idea of iron being the
50 dominant component of the core gained firm support from cosmochemical
51 and geochemical observations, seismic data, theory of geomagnetism, and
52 high-pressure studies. Cosmochemical data and studies of iron meteorites
53 give evidence that the Earth's core contains significant (5–15%) amounts of
54 nickel. Already in the early 1950s it was recognized that the outer core is
55 less dense than iron or an iron–nickel alloy at corresponding conditions.
56 Current estimates for the density deficit relative to solid iron vary between
57 6% and 10% for the outer core and 2–5% for inner core (*Dubrovinsky et al.*,
58 2000, 2007; *Dewale et al.*, 2006). Uncertainties are in part related to the
59 experimental difficulties in measurements of the properties of iron (or iron–
60 nickel alloys) at multimegabar pressure range and high temperatures, but
61 the fundamental problem is the absence of a universal pressure scale that
62 has led to significant discrepancies in the inter-laboratory studies (see, for
63 example, *Dubrovinsky et al.*, 2007). Nevertheless, the core is expected
64 to contain other element(s) apart from iron and nickel. In order to be
65 considered as an important component of Earth's outer core, an element
66 lighter than iron must be geochemically abundant, alloying readily with
67 iron, and partitioning preferably to liquid rather than to a solid iron phase.
68 These constraints have resulted in S, O, Si, C, and H being the preferred
69 candidates, but practically all geochemically abundant elements which are
70 lighter than iron were suggested and have their proponents. Additional
71 constrains come up from the analysis of elements depletion in the mantle
72 (*e.g.* how much of a certain element was available to be incorporated into the
73 bulk core in the past (on stage of the Earth differentiation) and at present),
74 the comparison of physical properties (density, compressibility, sound
75 velocities, *etc.*) of Fe–(Ni)-light element compounds with corresponding
76 properties of the core, the observation of the partitioning of elements at the
77 core–mantle and inner-outer core boundaries. Available information suggests
78 that none of elements alone could satisfy all constrains simultaneously.
79 Studies of elements solubility in molten iron in laser-heated diamond anvil
80 cells at megabar pressure range and temperatures in excess of 3,000 K by
81 means of analytical transmission electron microscopy (*Takafuji et al.*, 2005)
82 and measurements of compressional sound velocities in light element alloys
83 of iron at high-pressure by inelastic X-ray scattering (*Badro et al.*, 2007) gave
84 consistent results. The preferred model derived from these experimental
85 works is the outer core containing about 5.3 wt.% oxygen and 2.8 wt.%
86 silicon, and the inner core with about 2.3 wt.% silicon. Note that these

87 conclusions depend on the structure of the iron–nickel–light element(s)
 88 alloys at core conditions and on the assumed temperature of the Earth’s core.

89 3. Melting of Iron and Temperature of the Earth’ Core

90 One of the major uncertainties in the modern geophysics is temperature of
 91 the Earth’s core. This is a basic parameter, essential for modeling thermal
 92 budget and dynamics of the planet, including such surface processes as
 93 volcanism and tectonics. There is no way to obtain temperature distribution
 94 in the Earth core from direct observations. However, temperature of solidi-
 95 fication of the core material at inner-outer core boundary (ICB, ~330 GPa)
 96 could provide a good estimate of temperature at the depth of ~5,100 km,
 97 and melting temperature of core-forming material would give low bound of
 98 the temperature at core–mantle boundary (CMB, depth ~2,900 km and
 99 pressure ~136 GPa). Assuming that iron is the main component of the
 100 Earth’s core, necessary information could be extracted from an experimentally
 101 measured melting curve of iron at multimegabar pressures (Figure 2). At
 102 ICB conditions the only experiments possible at the moment are those
 103 based on shock waves (see, for example, *Nguyen and Holmes, 2004*). In these
 104 experiments temperature estimations are difficult and require significant
 105 assumptions. Studies in laser-heated diamond anvil cells (DACs) could be

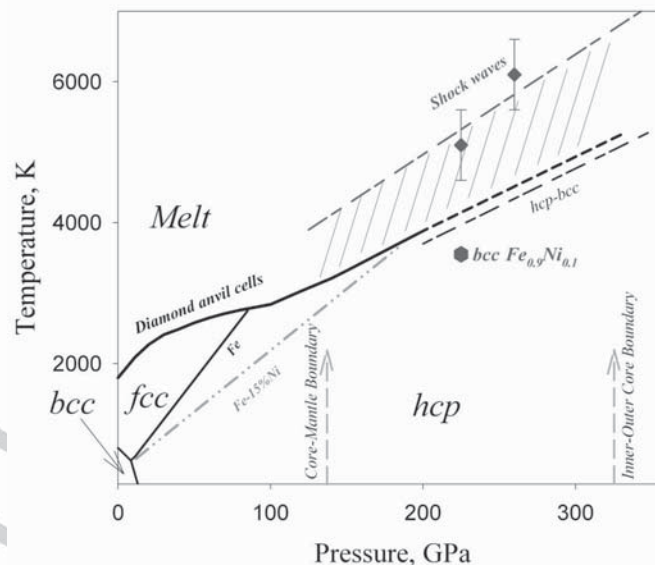


Figure 2. Representative phase diagram of iron and iron–nickel alloys at high pressures and temperatures. hcp iron is stable over a wide range of pressures and temperatures, while bcc iron is predicted to exist in the inner core (dashed line) and a bcc $\text{Fe}_{0.9}\text{Ni}_{0.1}$ alloy is experimentally observed at 225 GPa and 3,400 K (hexagon). Melting curves of iron from shockwave measured (diamonds) are much higher than static diamond cell results (dashed line). Shaded area: current survey of the melting temperatures of iron at core pressures.

106 extended to over 200 GPa, but a detection of the melting event in DACs is
107 not easy and even at a moderate pressure below 100 GPa there is no
108 unambiguous consensus. Currently the difference between the shock wave
109 and DAC results on iron melting at ICB is more than 1,000 K (Figure 2)
110 with different theoretical calculations supporting either of the estimates (*e.g.*
111 *Alfè et al.*, 2003; *Belonoshko et al.*, 2003). The effect of light elements on
112 melting temperature and sub-solidus relations in iron-based alloys at Earth's
113 core conditions adds further uncertainties in guessing temperature of the
114 Earth's core.

115 4. Structure of Iron and Iron-Based Alloys in the Earth's Inner Core

116 Iron has been reported to have several phases based on x-ray diffraction at
117 high pressure and temperature, *i.e.*, α and δ , body-centered-cubic (*bcc*); γ ,
118 face-centered-cubic (*fcc*); ϵ , hexagonal closed-packed (*hcp*); β , double-hcp
119 phase (*dhcp*) or orthorhombic phase. The α -, δ -, and γ -phases at lower
120 pressures are well established and broadly accepted. The ϵ -phase has
121 been proved to be a dominant phase in a wide pressure–temperature range
122 approaching the Earth's core conditions. The β -phase, observed at pressures
123 above 30 GPa and high temperature, is probably metastable, or stabilizes
124 due to contaminations.

125 While studying of pure iron at multimegabar pressures draw considerable
126 attention and have provided rich experimental data, the knowledge of the
127 behavior and properties of Fe–Ni alloys at conditions of the Earth's core is
128 still limited. At ambient pressure iron–nickel alloys with up to 25 at% of Ni
129 have the *bcc* structure, while higher nickel contents promote crystallization
130 of the *fcc*-structured phase (Figure 2). Compression of *bcc*-structured alloys
131 at ambient temperature results in their transformation to the *hcp*-phase at
132 pressures between 7 and 14 GPa (depending on composition and conditions
133 of experiments). Like in case of pure iron, no further transformations were
134 observed up to a pressure of ~ 300 GPa on compression of the Fe–Ni alloys
135 with up to 20 at% Ni. However, the presence of nickel significantly affects
136 phase relations in the Fe–Ni system at high temperatures and pressures.
137 While the slope of the *hcp*–*fcc* phase boundary for pure iron is 35–40
138 K/GPa, there are indications that the phase boundaries of Fe–Ni alloys with
139 10 to 30% Ni might have much lower slopes (15–25 K/GPa). Simple
140 extrapolation of experimental results (Figure 2) for Fe with 10–20 wt.% Ni
141 suggests that the stable phase at the inner core conditions is either hexagonal
142 close packed, or a mixture of *hcp* and *fcc* phases (*Lin et al.*, 2002; *Mao et al.*,
143 2006; *Dubrovinsky et al.*, 2007).

144 Unlike experiments, theoretical calculations do not require extrapolation
145 in modeling the state and properties of iron and iron-based alloys at
146 conditions of the Earth's core and suggest that at sufficiently high pressures
147 (above 100 GPa) and temperatures close to the melting point iron transforms
148 into *bcc* structured phase (*Belonoshko et al.*, 2003). Although for pure iron
149 this prediction has not been confirmed yet, the *hcp*–*bcc* transformation of the
150 $\text{Fe}_{0.9}\text{Ni}_{0.1}$ alloy was observed at pressures above 225 GPa and temperatures
151 over 3,400 K (*Dubrovinsky et al.*, 2007). Moreover, both experimental (*Lin*
152 *et al.*, 2003) and theoretical (*Vocadlo et al.*, 2003) works point out towards
153 stabilization of the *bcc*-structured phase due to the elements (particularly,
154 silicon) important for the Earth's inner core chemistry.

155 If *bcc*-structured phase is indeed dominant in the Earth's inner core,
156 it could help to clarify at least one of its enigmatic properties – elastic
157 anisotropy (Figure 1b). Preferred orientation of crystals of the *hcp*-structured
158 iron-based alloy could, in principle, explain why sound waves propagate
159 faster along Earth's spin axis than in the equatorial plane. However, elastic
160 anisotropy of *hcp* iron rapidly decreases with increasing temperature. In
161 contrast, according to molecular dynamic simulations (*Belonoshko et al.*,
162 2008) cubic *bcc* iron is extremely anisotropic and can account for 12% of
163 the seismic wave anisotropy sufficient for explanation of anisotropy of the
164 inner core. Simulations reveal also that abundant grain boundaries and
165 defects formed in *bcc* iron at high temperatures could lead to drastic
166 decrease of the shear modulus and shear wave velocity, as compared to
167 those estimates obtained from the averaged single-crystal values, thus
168 providing possible explanation of low rigidity of the Earth's inner core.

169 5. Conclusions

170 The emerging picture of the Earth's core shows that it as a very dynamic
171 region. Functioning of the geodynamo requires the outer core liquid with
172 small enough viscosity to permit fluid motion with typical velocities of
173 20 km/year. Intensive convection in the outer core is powered by secular
174 cooling which results in solidification at the inner core boundary and
175 growth of the inner core. Preferential fractionation of elements during
176 crystallization of the inner core changes composition, density, and chemical
177 activity of outer core. Continuous changes of the outer core provoke its
178 continuous reactions with the mantle and induce small-scale dynamic
179 processes at the core–mantle boundary. Mineral physics plays a key role in
180 providing data about the state and properties of the materials at conditions
181 of the Earth's core. The main challenge of the studies of mineralogy of the
182 Earth's core is to find a common ground for explanation of different,

183 sometimes seemingly unconnected or controversial, properties – the chemical
184 and phase composition, temperature and the heat flow, nature of the dynamics
185 and geomagnetism, convection and elasticity.

186 **References**

- 187 Alfè, D., M. J. Gillan, G. D. Price (2003) Thermodynamics from first principles: temperature
188 and composition of the Earth's core. *Mineral. Mag.*, 67, 113–123.
- 189 Antonangeli, D., Occelli, F., Requardt, H., Badro, J., Fiquet, F., Krisch, M., et al. (2004)
190 Elastic anisotropy in textured hcp-iron to 112 GPa from sound wave propagation
191 measurements. *Earth Planet. Sci. Lett.*, 225, 243–251.
- 192 Badro, J., et al. (2007) Effect of light elements on the sound velocities in solid iron:
193 Implications for the composition of Earth's core. *Earth Planet. Sci. Lett.*, 254, 233–238.
- 194 Belonoshko, A. B., R. Ahuja, B. Johansson, (2003) Stability of the body-centred-cubic phase
195 of iron in the Earth's inner core. *Nature*, 424, 1032–1034.
- 196 Belonoshko, A. B., et al. (2008), Elastic anisotropy of Earth's inner core. *Science*, 319, 797–
197 800.
- 198 Dubrovinsky, L., et al. (2000) X-ray study of thermal expansion and phase transition of iron
199 at multimegabar pressure. *Phys. Rev. Lett.*, 84, 1720–1723.
- 200 Dubrovinsky, L., et al. (2007) Body-centered cubic iron-nickel alloy in Earth's core. *Science*,
201 316, 1880–1883.
- 202 Dewaele, A., et al., (2006) Quasihydrostatic equation of state of iron above 2 Mbar. *Phys.*
203 *Rev. Lett.*, 97, 215504.
- 204 Lin, J.-F. et al. (2002) Iron-nickel alloy in the Earth's core. *Geophys. Res. Lett.*, 29, 1471,
205 doi:10.1029/2002GL015089.
- 206 Lin, J.-F., et al. (2003) Iron-silicon alloy in Earth's core? *Science*, 295, 313–315.
- 207 Mao, W. L., et al. (2006) Phase relations of Fe-Ni alloys at high pressure and temperature.
208 *Phys. Earth Planet. Interiors*, 155, 146–150.
- 209 Nguyen, J. H., N. C. Homes (2004) Melting of iron at the physical conditions of the Earth's
210 core. *Nature*, 427, 339–341.
- 211 Takafuji, N., et al. (2005) Solubilities of O and Si in liquid iron in equilibrium with
212 (Mg,Fe)SiO₃ perovskite and the light elements in the core. *Geophys. Res. Lett.*, 32,
213 L06313, doi:10.1029/2005GL022773.
- 214 Vocadlo, L., et al. (2003) Possible thermal and chemical stabilization of body-centred-cubic
215 iron in the Earth's core. *Nature*, 424, 536–539.